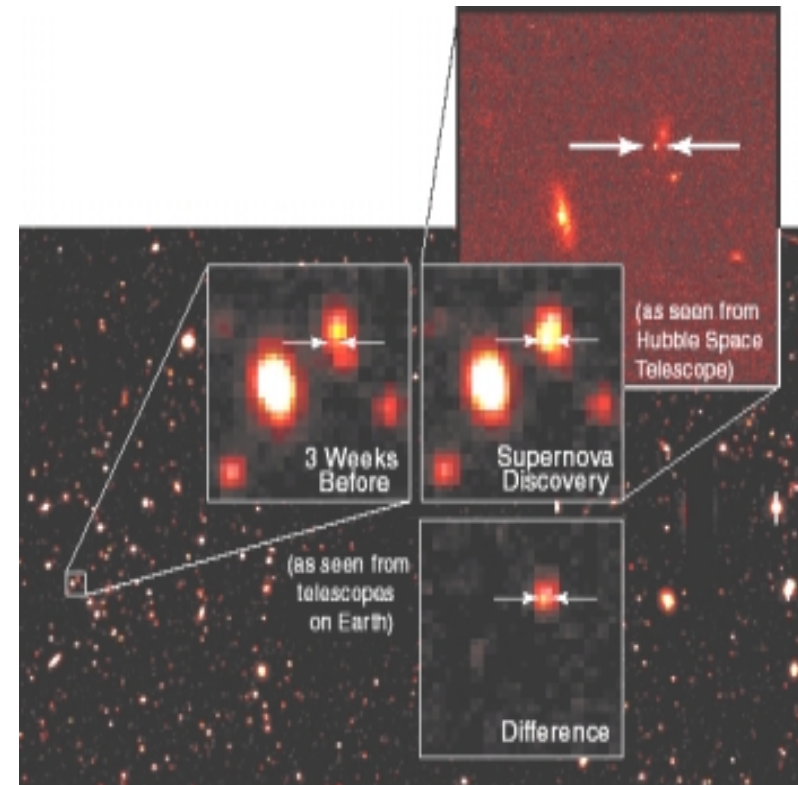
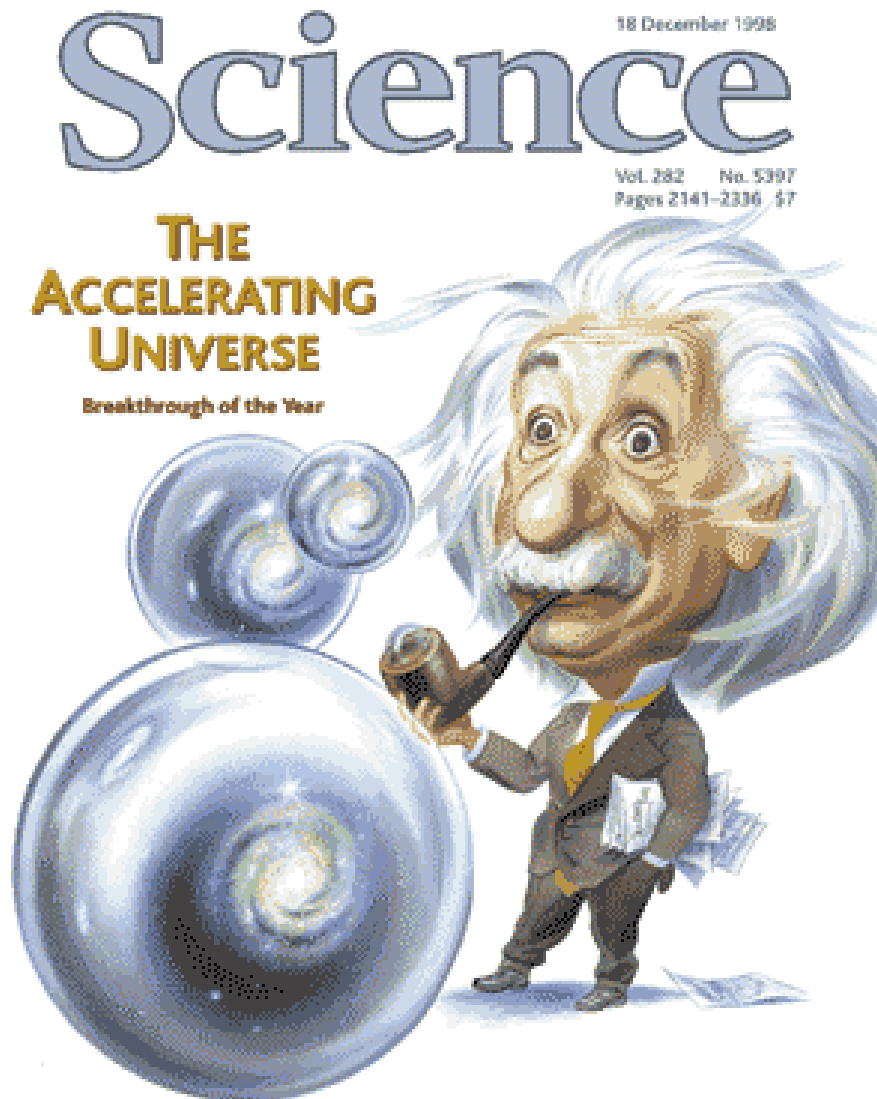


# Astrophysics to Understand the Universe

## Mass Density, Vacuum Energy Density, and Curvature



## Fundamental Questions:

- *Will the universe last forever?*
- *Is the universe infinite?*
- *What is the universe made of?*

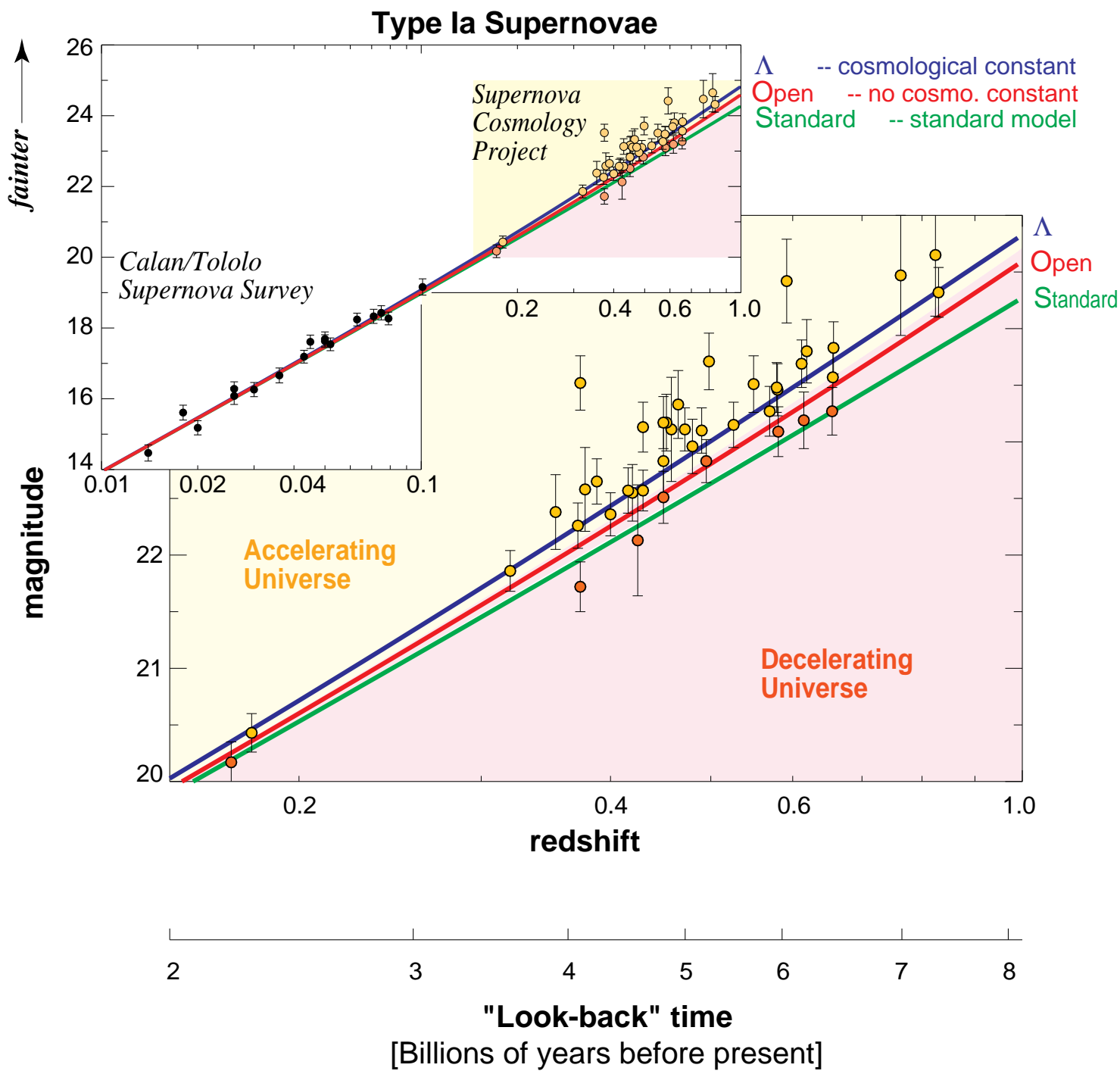
## An unusual moment in human history:

At the beginning of this century, Einstein developed the conceptual tools to address these questions empirically.

In the past decade or so, technology has advanced to the point that we can now make the measurements that begin to answer these fundamental questions.

Progress is now being made with large scientific programs, including the Supernova Cosmology Project and the Cosmic Microwave Background satellites: MAP, and PLANCK.

---



The implications of an accelerating universe:

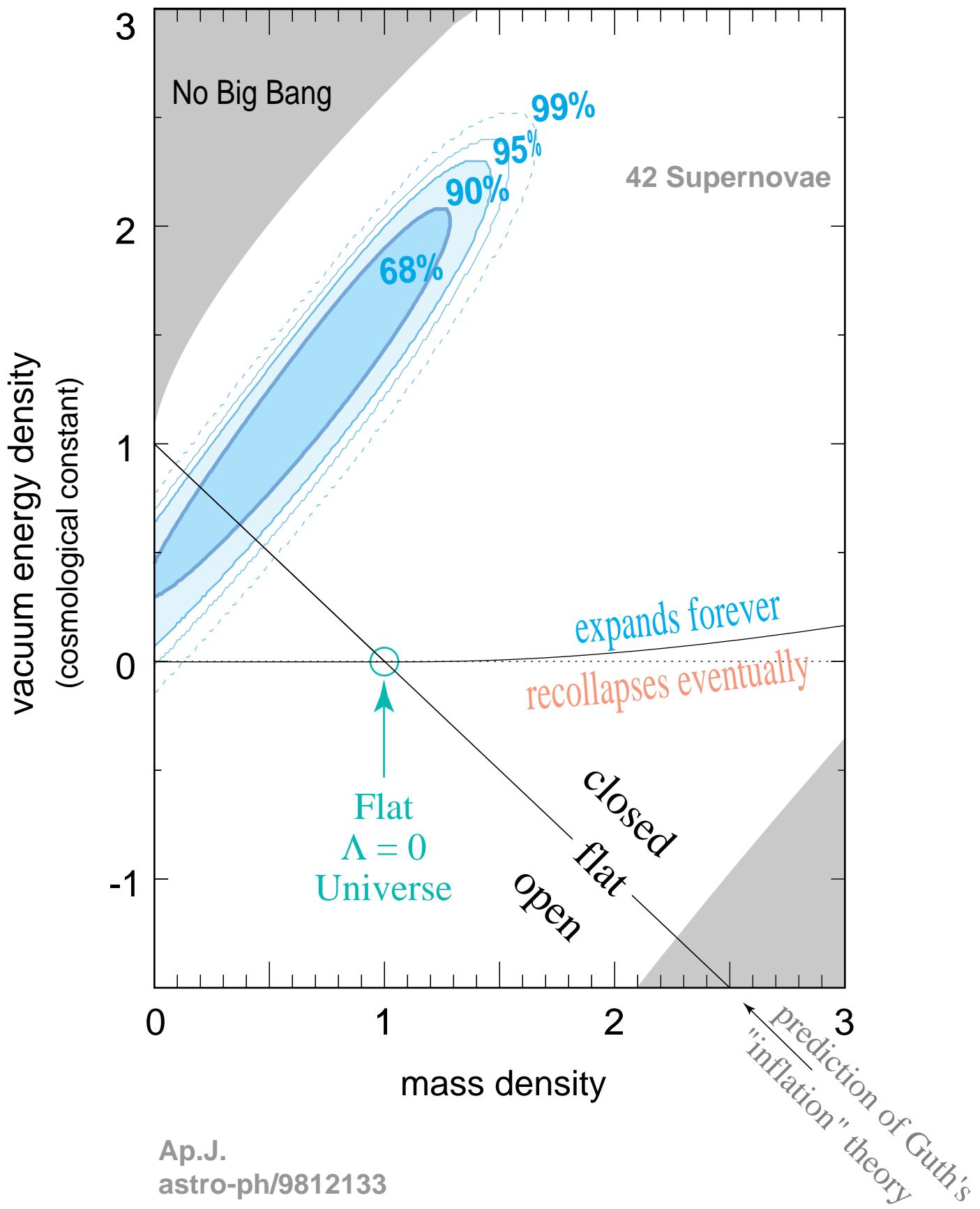
1. The expansion is not slowing to a halt and then collapsing (i.e., the universe is *not* "coming to an end").  
In the simplest models, it will expand forever.
2. There is a previously unseen energy pervading all of space that accelerates the universe's expansion.

This new accelerating energy ("dark energy") has a larger energy density than the mass density of the universe (or else the universe's expansion wouldn't be accelerating).

What we don't know is:

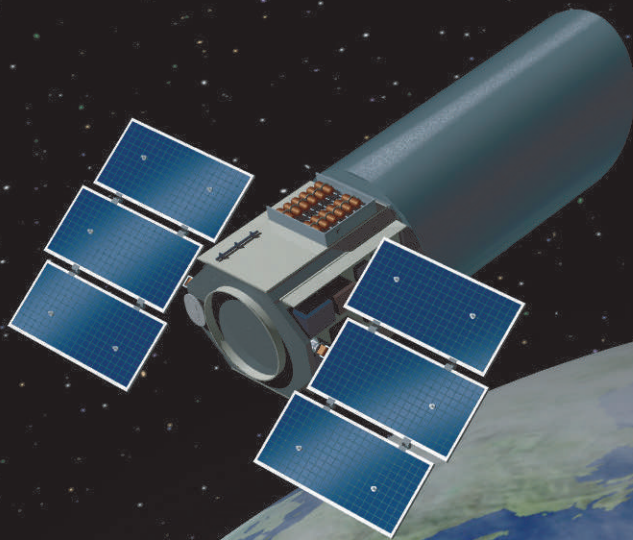
1. How much of mass density and dark energy density is there? I.e., how much dark matter and dark energy do we need to look for?  
The answer to this question determines the "curvature" of the universe, and can tell us about the extent of the universe: infinite or finite.
2. What is the "dark energy"? Particle physics theory proposes a number of alternatives, each with different properties that we can measure. Each of the alternative theories raises some important questions/problems of fundamental physics.

Supernova Cosmology Project  
Perlmutter *et al.* (1998)





SNAP: Supernova / Acceleration Probe





supernova / acceleration probe

***satellite overview***

- **1.8m aperture telescope**

*Can reach very distant SNe.*

- **1 square degree mosaic camera, 1 billion pixels**

*Efficiently studies large numbers of SNe.*

- **3-channel spectroscopy, 0.3um -- 1.8um**

*Detailed analysis of each SN.*

*MIDEX+ class satellite:*

Dedicated instrument.

Designed to repeatedly observe an area of sky.

Essentially no moving parts.

4-year construction cycle.

3-year operation for experiment  
(lifetime open-ended).



supernova acceleration probe  
***satellite baseline specifications***

<b><i>Aperture</i></b>	1.8m
<b><i>F.O.V.</i></b>	1 square degree
<b><i>Launch Vehicle</i></b>	Delta II 7920

<b><i>Image array</i></b>	36k x 36k mosaic CCD
<b><i>Architecture</i></b>	2k x 4k 10-micron pixels
<b><i>Plate Scale</i></b>	10 micron/0.10 arc-sec

<b><i>Image Array Bands</i></b>	B,V,R,I,Z
<b><i>Q.E.</i></b>	65% @ 1000nm (LBNL CCD) 92% @ 900nm >85% @ 400nm-800nm
<b><i>Read Noise</i></b>	<2-4 e-
<b><i>Dark Current</i></b>	0.08 e-/min/pixel
<b><i>Image Array Temp.</i></b>	150K

<b><i>Spectroscopy</i></b>	3-channel pick-off: 325-600nm "Blue" CCD + 600-1050nm "Red" LBL CCD + 1050-1800nm IR (HgCdTe)
<b><i>Thermal Cooling</i></b>	passive

<b><i>Data Downlink</i></b>	320 Mbps
<b><i>Data Storage</i></b>	256 Gbits solid state

<b><i>Pointing Stability</i></b>	<0.1 arc-sec
<b><i>Image Stabilization</i></b>	2-axis fast steering mirror
<b><i>Satellite Mass</i></b>	<1500kg
<b><i>Power Consumption</i></b>	<800W



# SNAP SAT

supernova acceleration probe  
***one-year baseline data package***

## Full sample of 2000 SNe between $z = 0.3$ and $1.7$

Discovery within  $\sim 2$  days of explosion (i.e.  $\sim 2$  weeks before max).  
Most dense coverage between  $z = 0.3$  and  $1.0$

- Spectra at max for all SNe. ( $0.3 - 1.8\mu\text{m}$ )
- Lightcurve points at least 1/week (restframe)  
from  $-15$  to  $+60$  days (restframe)

## Discover every SN in the field

Nearly continuous monitoring of  
     $\sim 2$  sq. deg. to  $m_{AB} (@ 1\mu\text{m}) \approx 28.5+$   
     $\sim 10$  sq. deg. to  $m_{AB} (@ 1\mu\text{m}) \approx 25+$   
     $\sim 100$  sq. deg. to  $m_{AB} (@ 1\mu\text{m}) \approx 24+$

## Satellite follows additional 200 SNe at $z < 0.15$

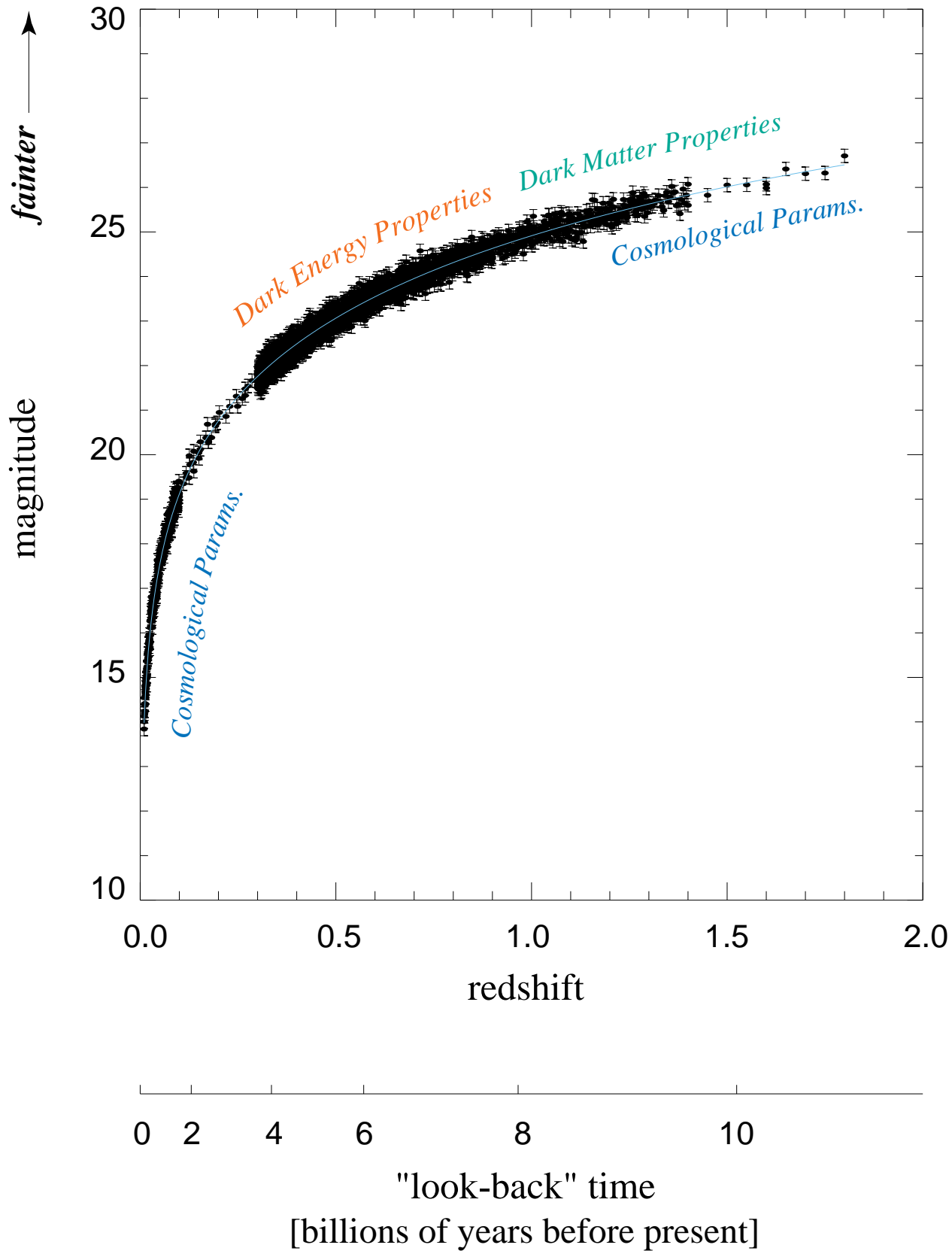
Discovery from ground, follow-up from space to ensure matching  
spectroscopy/photometry across redshifts.

## Subsample of 200 SNe from the full sample

- Selected to span
- lightcurve timescales
  - galaxy environments  
(morphology, galactocentric radius)
  - redshifts
- 
- Spectra at least 2/week (restframe) first month  
1/week (restframe) later
  - Synthetic "filter-tuned" photometry from spectra  
for perfect K-corrections

Baseline One-Year Sample  
2000 SNe

**SNAP** Dark Energy Observer





supernova acceleration probe  
***baseline science goals***

<i>Target Parameter</i>	<i>Constraint</i>	<i>Target Statistical Uncertainty</i>
-----------------------------	-------------------	---

**Determine mass density, vacuum energy density, and curvature**

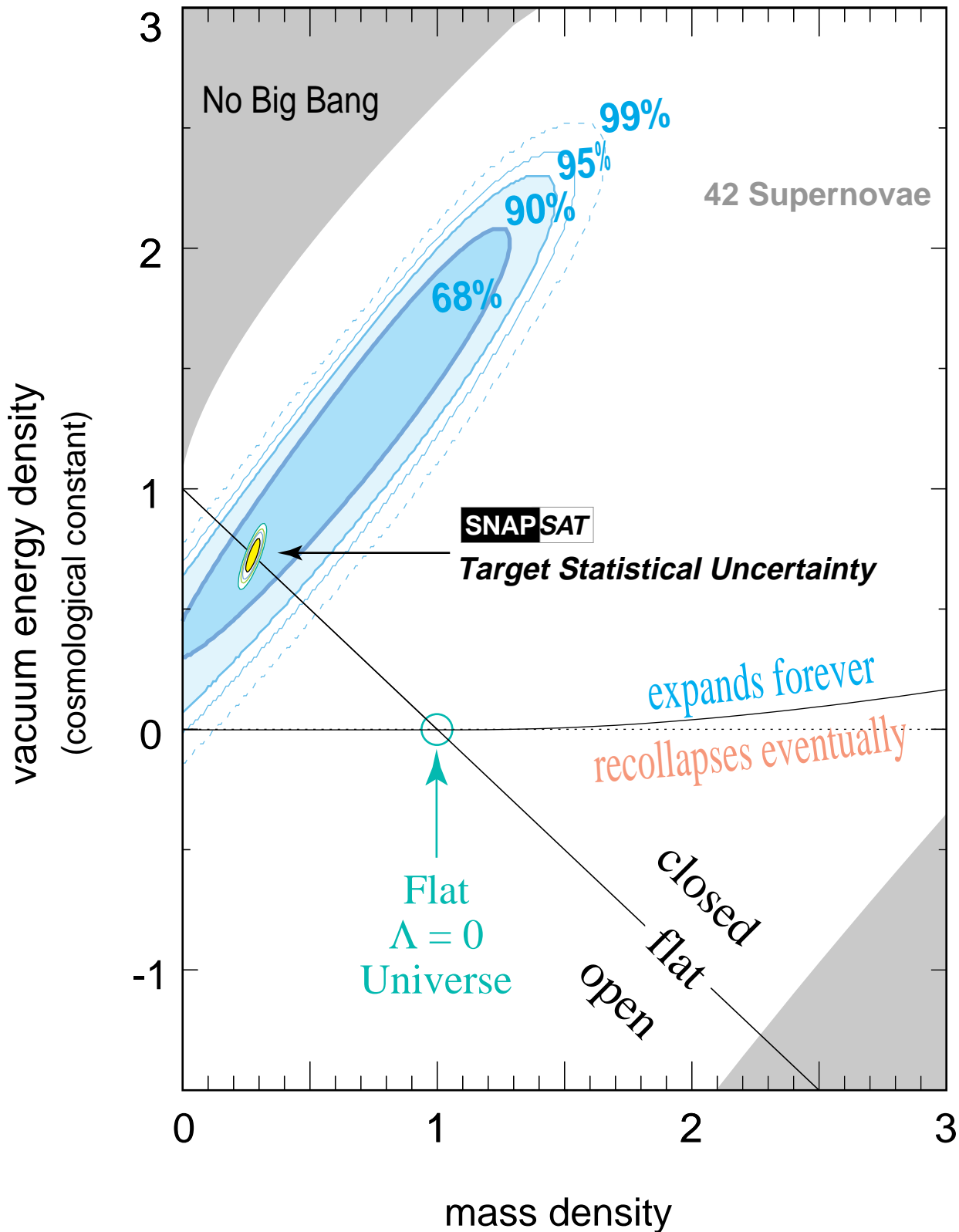
mass/vacuum energy $\Omega_M, \Omega_\Lambda$	in flat universe $\Omega_k=0$	0.01
curvature $\Omega_k$	(indep. of CMB)	0.05
mass density $\Omega_M$	unconstrained	0.02
vacuum energy $\Omega_\Lambda$	"	0.05

**Properties of Dark Energy**

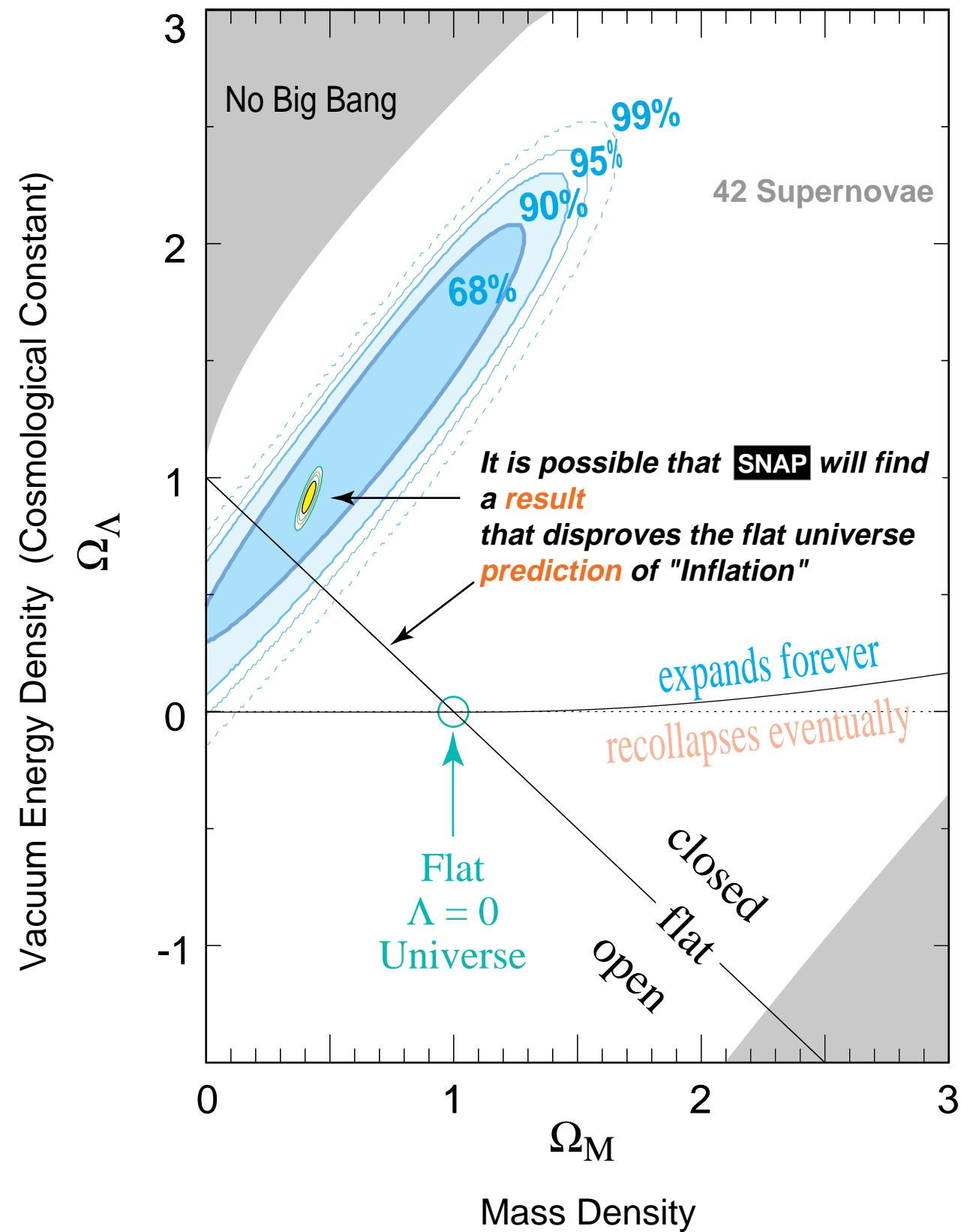
Eq. of State $w$ vs $\Omega_M$	in flat universe	
Eq. of State $w$	with $\Omega_M \approx 0.3$	0.05
Study time-varying $w(t)$ by studying $d_L(z)$ with $\Delta z \approx 0.03$ bins.		

***A definitive supernova cosmology measurement.***

Supernova Cosmology Project  
Perlmutter *et al.* (1998)



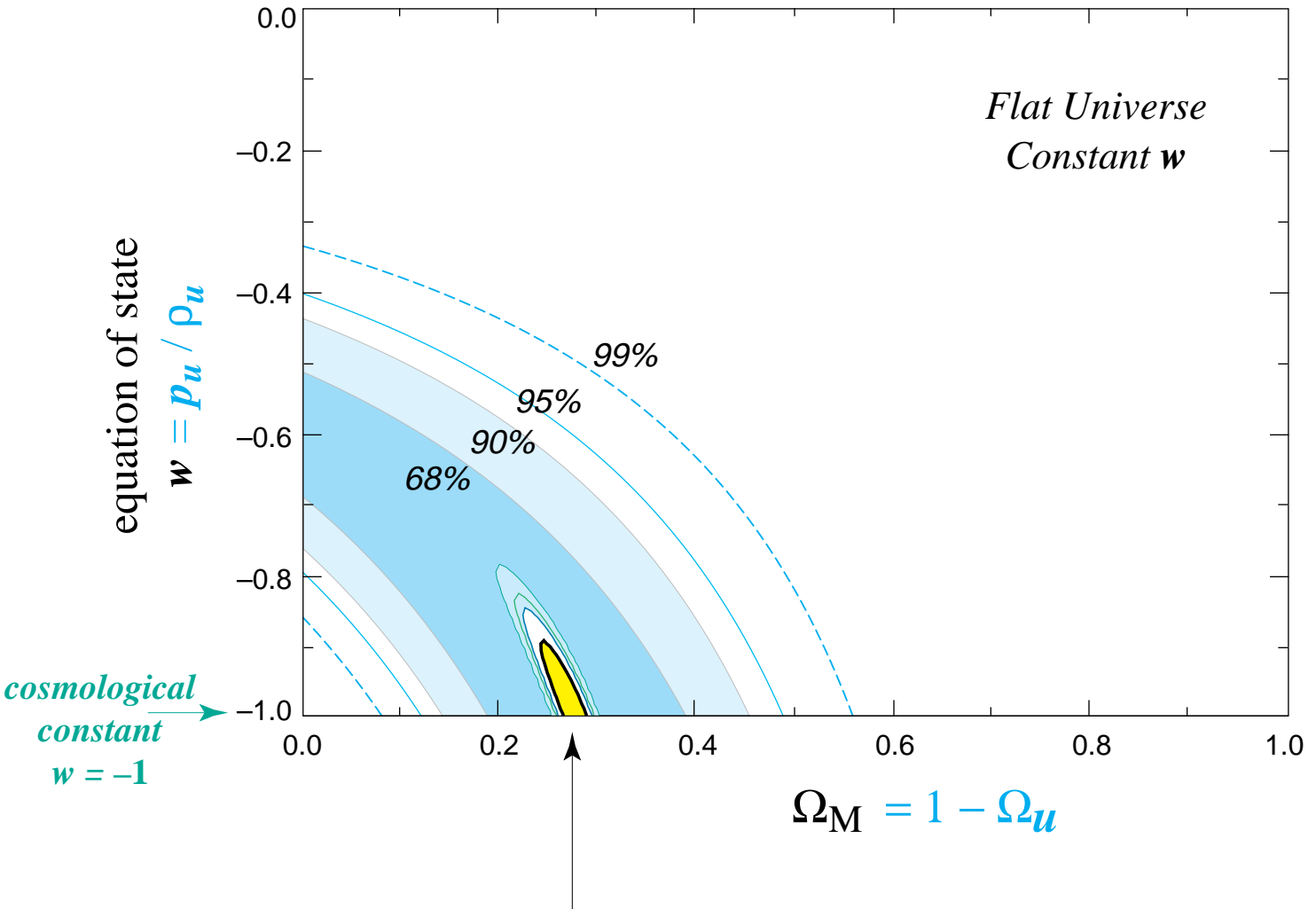
Supernova Cosmology Project  
Perlmutter *et al.* (1998)



# Dark Energy

Unknown Component,  $\Omega_u$ , of Energy Density

Supernova Cosmology Project  
Perlmutter *et al.* (1998)

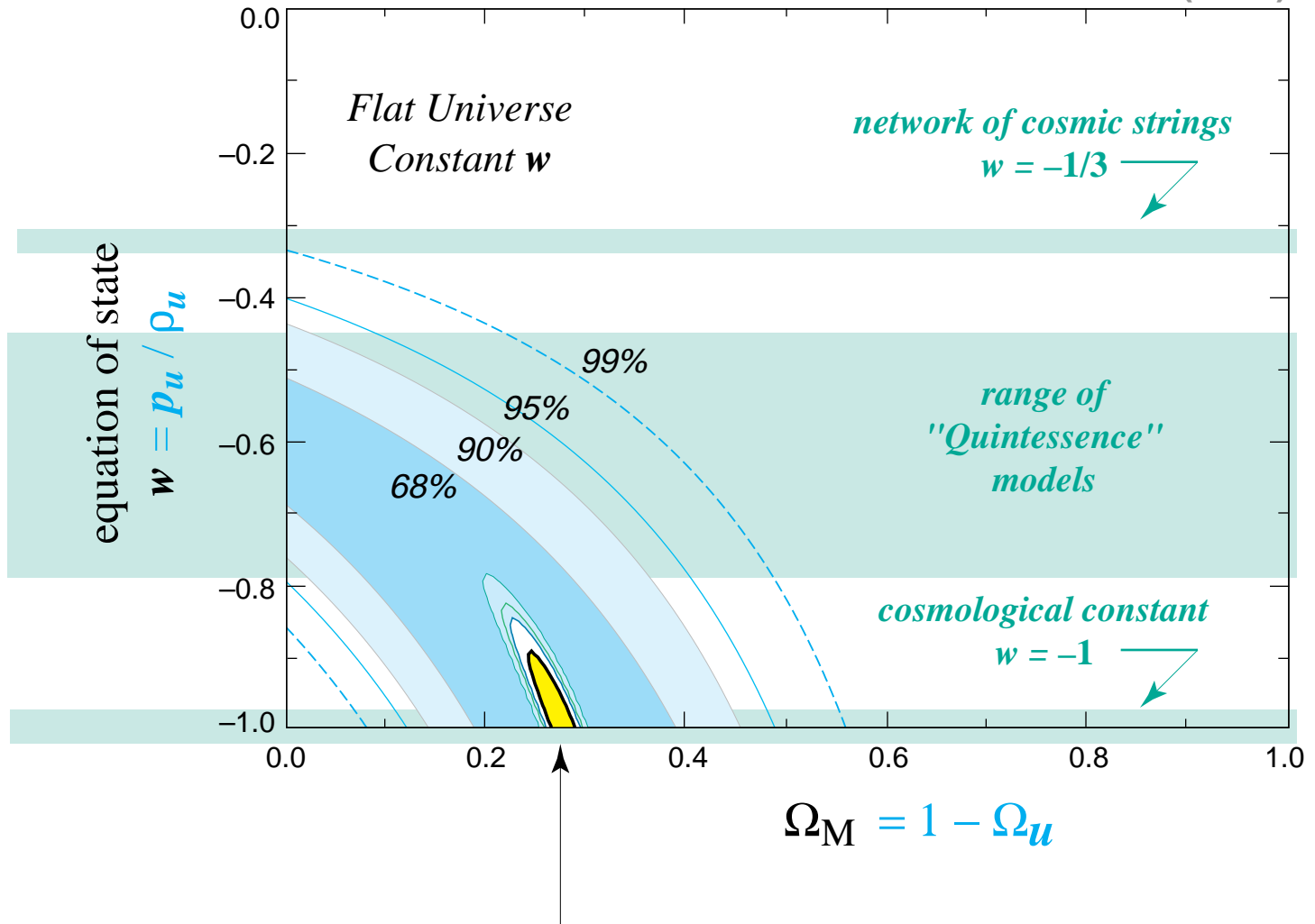


**SnapSat**  
**Target Statistical Uncertainty**

# Dark Energy

Unknown Component,  $\Omega_u$ , of Energy Density

Supernova Cosmology Project  
Perlmutter *et al.* (1998)



**SNAP Satellite  
Target Statistical Uncertainty**

# SNAP SAT

supernova acceleration probe  
***controls for systematics***

## Dust and Extinction

Determine individual extinctions laws for a given supernova,  
using range of color measurements into near IR.

Determine gray dust contribution (if any), tracking  $d_L(z)$  to high  $z$ .

## SN Progenitor Age and Metallicity Effects

Studies of lightcurve-timescale (and spectral) luminosity indicators  
with significant sample of SNe in all host galaxy environments.

## Measurement Systematics

Observe all SNe at all redshifts with single, calibrated, stabile  
photometry/spectroscopy system.  
(Avoid multiple instruments with different atmospheric/moon  
conditions.)





supernova acceleration probe  
***complementary science***

### Cosmological Parameters...

Type II supernova expanding photosphere,  $\Omega_M$ ,  $\Omega_\Lambda$  at  $z \approx 0.5$   
Strong lensing statistics.  $\Omega_\Lambda$   
Weak lensing  
Galaxy clustering,  $P(k)$   
 $z > 1$  clusters and associated lensing  
...

### ...and Beyond

GRB optical counterparts: rates, lightcurves, and spectra  
MACHO optical counterparts by proper motion  
Galaxy populations and morphology to co-added  $m \approx 32$   
Target selection for NGST  
Kuiper belt objects  
Supernova rates, star formation rates  
Supernova phenomenology studies  
Low surface brightness galaxies, luminosity function  
...

# Roadmap for the Supernova Cosmology Project

	Nearby SN Factory	Subaru +Keck+HST Key Project	NGST Key Project	SNAPSat
SNe / Year				
@ $z < 0.06$	100	.	.	100
@ $z \sim 0.5$	.	12	.	1000
@ $z > 1$	.	4	20	900
Duration	2 yrs	3 yrs	3 yrs	2 yrs
Calibrated Candle	✓	✓	✓	✓
Reddening Measure	✓	✓	✓	✓
Gray Dust Measure	✓	✓	✓	✓
Metallicity Measure	.	.	✓	✓
Evolution Checks	✓	.	.	✓
Standardized Datasets	.	.	.	✓

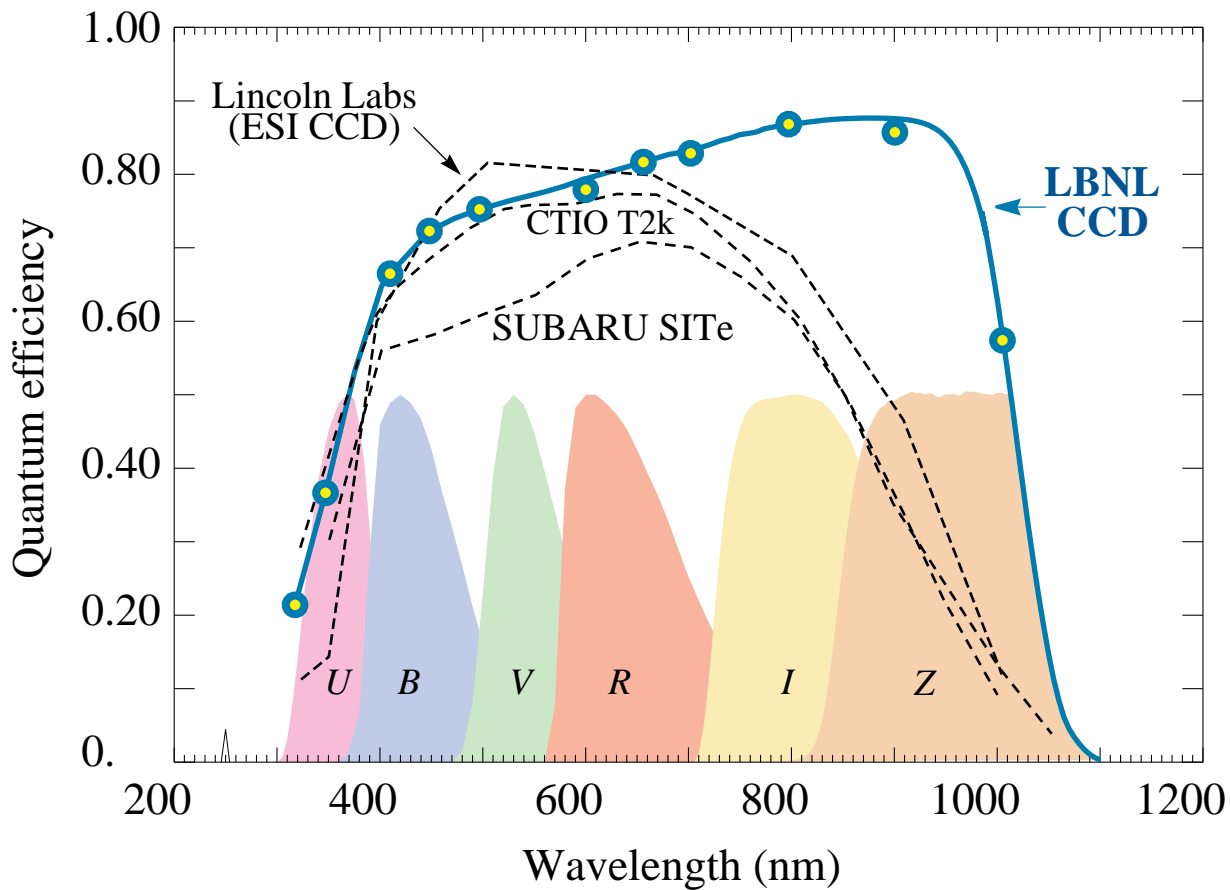
# LBLN CCD Technology

High quantum efficiency from near UV to near IR

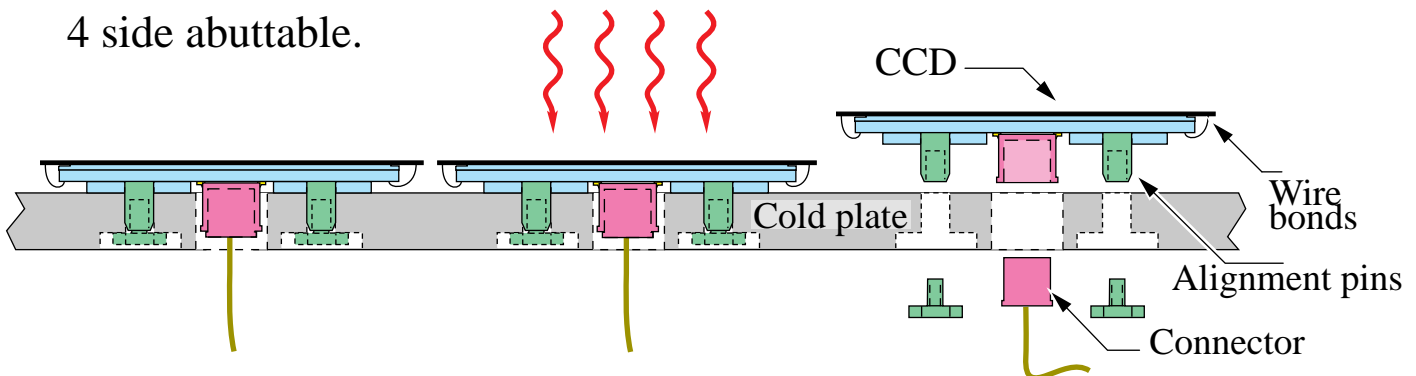
No thinning, no fringing.

High yield.

Radiation hard.



4 side abutable.



# NASA Interest in New Astrophysics Probes

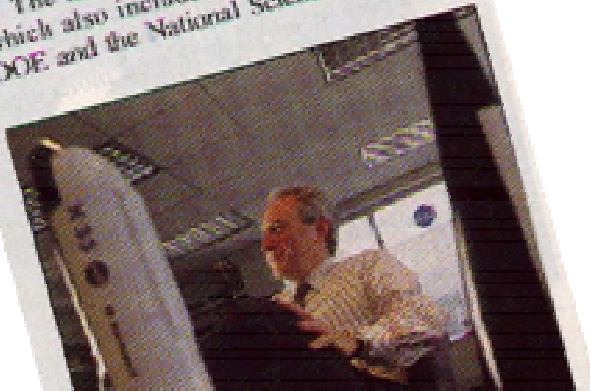


## PHYSICS

### Come Fly With Me, Goldin Tells Physicists

BATAVIA, ILLINOIS—Space is the final frontier for particle physics, NASA Administrator Daniel Goldin declared in a 28 May press conference here at the Fermi National Accelerator Laboratory (Fermilab). But Goldin's vision of joining forces with the Department of Energy (DOE) and other agencies in an all-out assault on the mysteries of gravity and high-energy physics failed to uplift some listeners when he lauded Earth-bound accelerators—the focus of DOE's high-energy physics program—a “smokestack approach” to research.

The message of the press conference, which also included representatives from DOE and the National Science Foundation



“... in the quest to quantify the expansion of the Universe, astronomers may have uncovered a new physical phenomenon -- a kind of vacuum energy they call the Cosmological Constant. This exciting discovery may have given astronomers and physicists an unexpected clue about fundamental physics.”

Daniel S. Goldin, May 28, 1999

Fermilab Inner Space-Outer Space Conference